

A TRIBOELECTRIC SENSOR ARRAY FOR WINDBORNE MARTIAN DUST

C.I. Calle¹, J.G. Mantovani², E.E. Groop¹, C.R. Buhler³, M.D. Hogue¹, A.W. Nowicki⁴

¹Electromagnetic Physics Laboratory, NASA Kennedy Space Center

²Florida Institute of Technology, Melbourne, Florida

³ Swales Aerospace, Kennedy Space Center

⁴Dynacs Inc., Kennedy Space Center

European Geophysical Conference XXVII, Nice, France, 21-26 April 2002



Introduction

 Triboelectric Sensor Array was developed to measure electrostatic interaction between windborne dust and several polymers

 Parent Technology: MECA Electrometer for 2001 Mars Odyssey cancelled lander

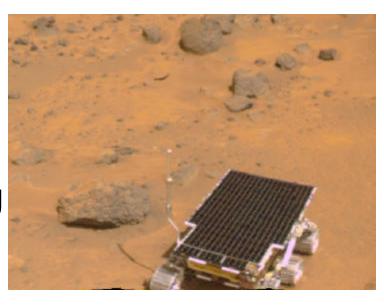




Martian Dust Environment

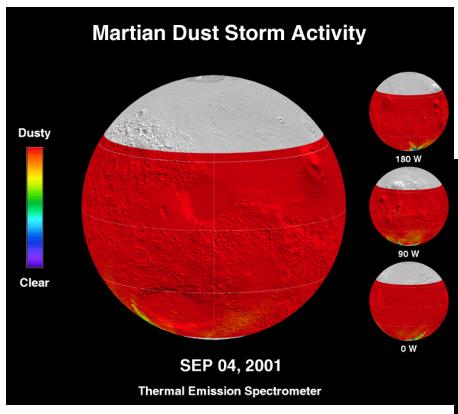
- Average dust particle in the Martian atmosphere: ~1 ?m in diameter
- Incident UV radiation may charge surface soil and dust particles on Mars
- Contact and frictional charge may also occur
- Future landing missions require a better understanding of electrostatic properties of dust and sand particles

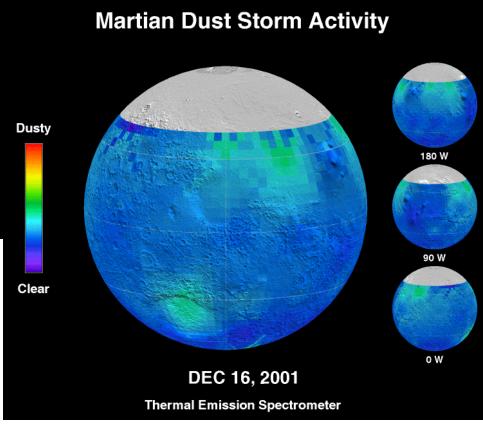






Martian Duststorms







Martian Dust Devils

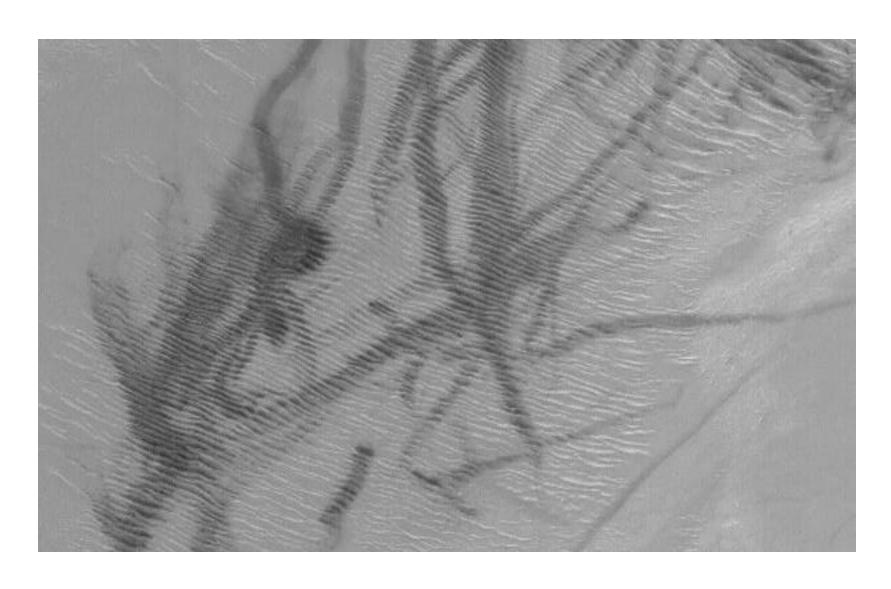
 Dust devils have been observed by the Mars Global Surveyor's Orbiter Camera (MOC) to be 2 km in diameter and 8 km in altitude



MGS/MOC May 13, 1999



Martian Dust Devils





Electrostatic Charging

- Incident UV radiation may charge surface soil and dust particles on Mars
- Contact and frictional charge may also occur
- Future landing missions require a better understanding of electrostatic properties of dust and sand particles

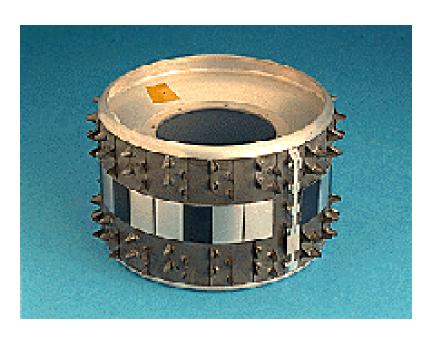


Martian Electrostatic Properties

- Most of what is known comes from earth-based measurements
 - Radar, radio occultation of spacecraft, microwave radiometry
 - Consistent with direct measurements of lunar rocks
 - Low conductivities



Martian In situ Experiment





- Wheel Abrasion Experiment (WAE) on Pathfinder used thin films of Al, Ni, and Pt, (200A -1000A), deposited on black, anodized Al strips attached to the rover wheel.
- As the wheel moved across the martian surface, a photovoltaic sensor was used to monitor changes in film reflectivity.
- Dust accumulation due to contact and frictional charging



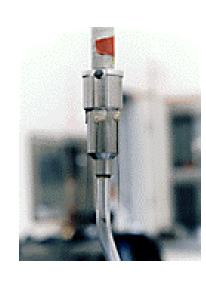
Ground Experiments

- 1973 lab experiments in Martian-like atmosphere:
 - Dust particle q? 10^4 e⁻¹
- In dusty, turbulent Martian environment:
 - -E? 5 kV/m



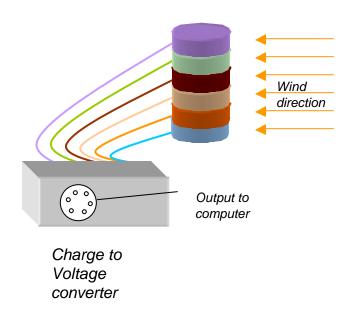
Pathfinder Rover

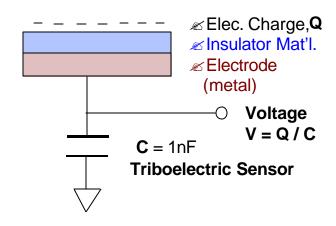
- Model of Sojourner wheel
- SME and simulant
- Potentials? 100 V
- Av arc times of 1?s
- /? 10 mA
- Discharge points to Sojuourner antenna base





Triboelectric Multi-sensor

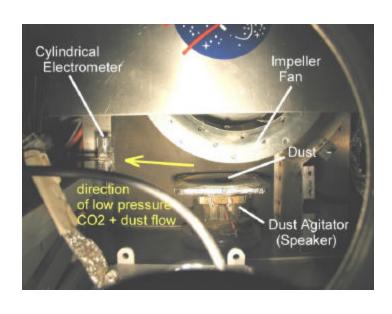




- Aerodynamic Multisensor Electrometer
- Expose different materials to "Martian" wind



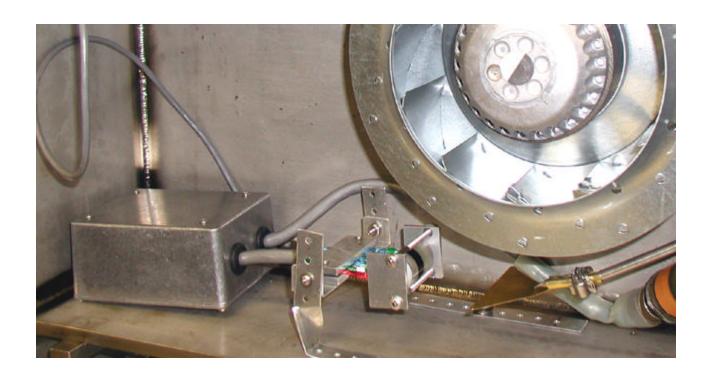
Dust Impeller



- Difficult to lift dust particles in the low atmospheric pressures near Martian surface
- Dust Impeller: vacuum chamber that uses vibrating membrane or feeder to deposit dust in front of rotating fan
- Speeds of 30 m/s



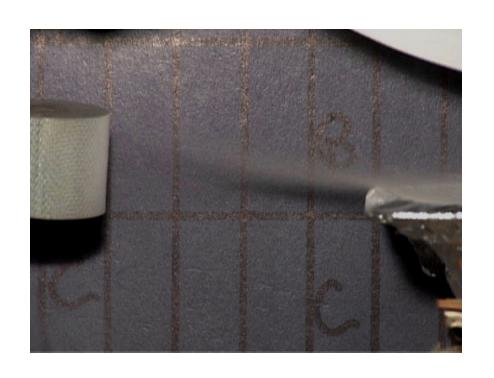
Multisensor/Dust Impeller

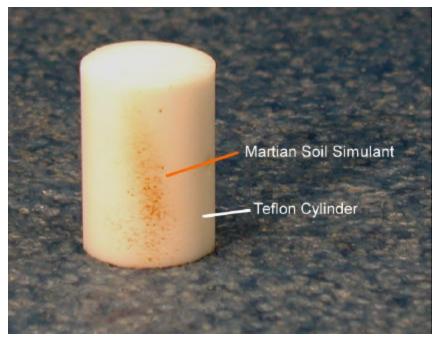


 Multisensor Electrometer in the Dust Impeller chamber.



Wind Simulation Experiments



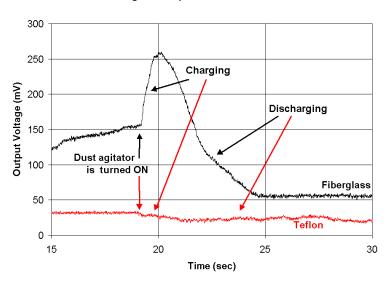


- Propel 5 to 20 ?m particles at samples under SME
- Q? -5 to + 19 pC



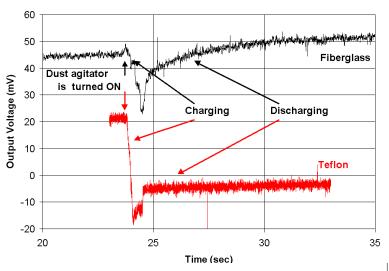
Experimental Results

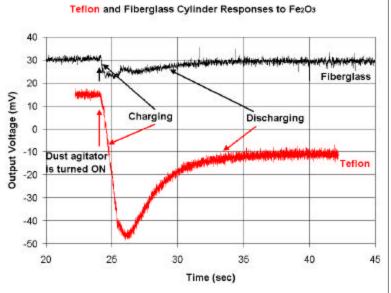
Teflon and Fiberglass Responses to Fine Martian Simulant



- JSC Mars-1 simulant, SiO₂, and Fe₂O₃ particles on Teflon and Fiberglass
- CO₂ at 9 mbars

Teflon and Fiberglass Cylinder Responses to Coarse SiO₂





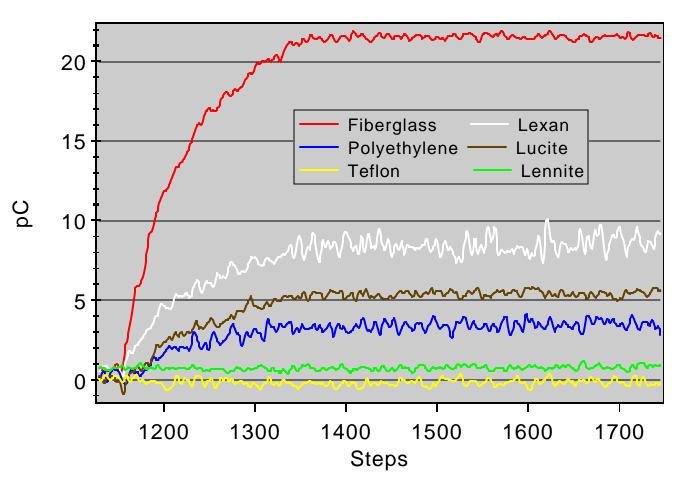


Triboelectric Series

Polymer	Mineral
Most Positive	
	Fe ₂ O ₃
Fiberglass	
Lucite	
	CaO
	Al_2O_3
	SiO ₂
	JSC Mars-1
Teflon	
Most Negative	



Results



- JSC Mars-1 simulant particles at 30 m/s
- CO₂ at 9 mbar



Conclusions

- Most of what we know about electrostatics on Mars: from ground based experiments
- No experiment has been flown designed solely for electrostatics
- We are currently developing instrumentation to study electrostatic properties of Martian dust
- Instruments will be proposed for future missions
- Experiments show that we can
- EMPL Website: http://empl.ksc.nasa.gov